

An Energy Balance Climate Model Which Accounts For Changes In Heat Storage Of The Earth

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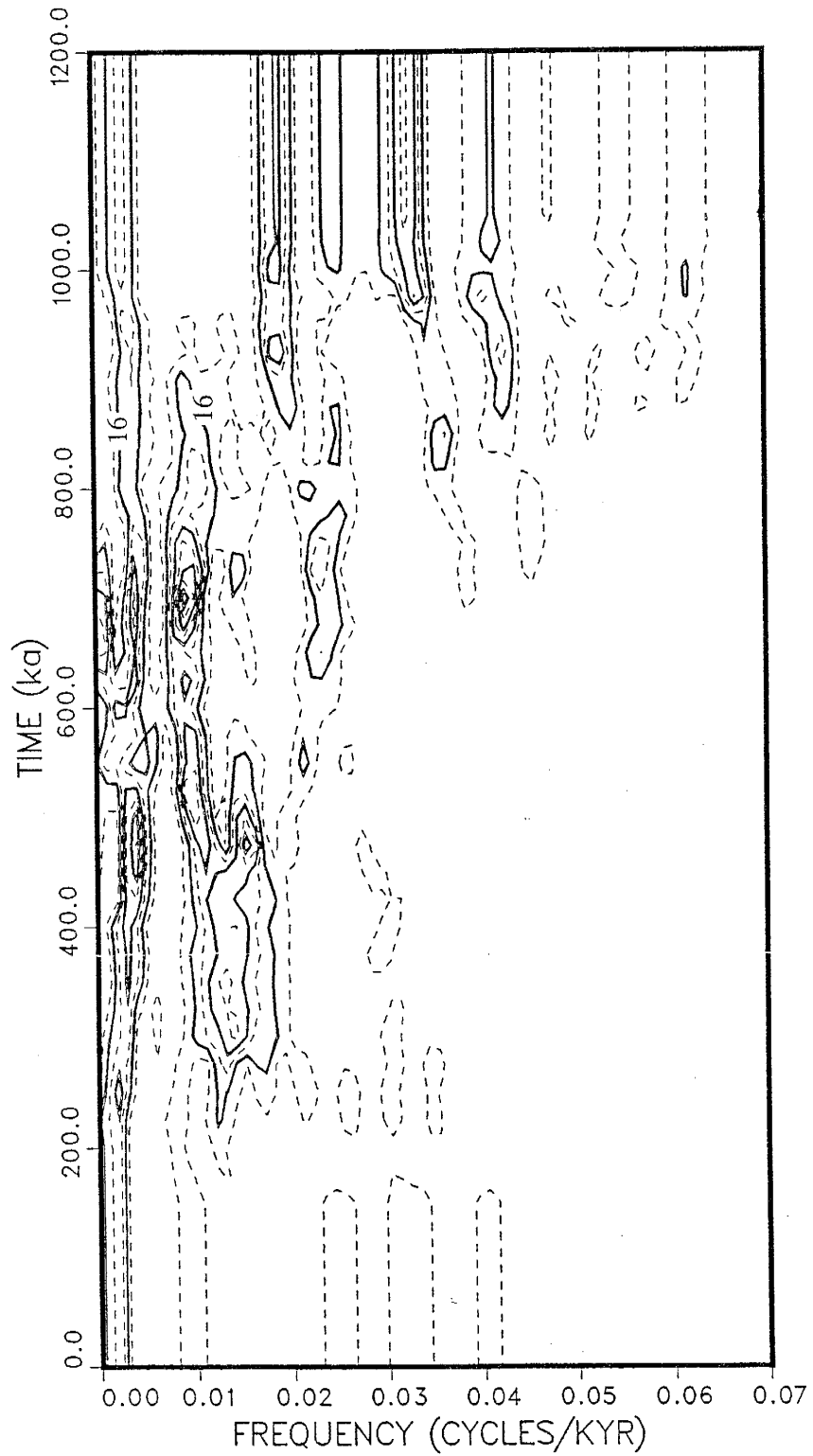
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Introduction

Energy balance climate models have proved a useful method for simulation and physical understanding of earth temperature distributions. Early nontransient models by Budyko (1969) and Sellers (1969) accurately reproduced current average yearly surface temperatures. Sellers (1973) obtained enviable results with a simple transient model run for a one year period. Short term transient energy balance models may be successful despite either ignoring change in heat storage of the earth or examining it only to a depth of a meter or less. For models simulating longer time periods it is subject to conjecture whether change in heat storage for the deep earth will significantly effect surface temperatures. The energy balance model presented here, while not designed to precisely imitate historical temperatures, demonstrates the relationship between deep earth heat storage and surface temperature oscillations.

Spectral analyses of past climate have demonstrated a pronounced one hundred year cycle. Fluctuations in the earth 's orbital parameters give rise to dominant frequencies of 19, 23, and 41 thousand years. For a climate which responds linearly with solar insolation variations, the one hundred thousand year cycle cannot be justified. Snieder (1985), however, has shown that in a nonlinear system the nonlinearity can give rise to strong one hundred thousand year component. Phillips et al. (1990) used water and isotopic mass balance in a transient numerical energy balance model to simulate 1.2 Million years of lake level history at Searles Lake, California. A spectral analysis of his results (figure 1) reveals periodicities which cannot be accounted for by simple linear solar insolation forcing, in particular, a strong 400 kyr component. If changes in deep earth heat storage proves to be a substantial contributor of nonlinear feedbacks in the energy budget, a comparison with this established reconstruction of climate history would be prompted.

figure 1.
SPECTRAL DENSITY CONTOUR PLOT , SEARLES LAKE



The Model

Energy Balance

The energy balance equation used for the model is

$$R = A + G_E \quad (1)$$

where,

R : Net radiation at surface of earth [kcal/cm² month]

A : Gain or loss of heat as a result of the atmosphere and hydrosphere circulation [kcal/cm² month]. Negative for gain of heat.

G_E: Change in heat storage of subsurface earth [kcal/cm² month]. Negative for loss of heat from the earth.

The net radiation term is expanded to its incoming and outgoing terms

$$R = Q (1 - \alpha) - I$$

where,

Q : solar radiation coming to the outer boundary of the atmosphere [kcal/cm² month]

α : albedo [-]

I : outgoing radiation [kcal/cm² month] .

The circulation term, A, and the outgoing radiation term, I, are related to a given latitudinal surface temperature using empirical relationships presented by Budyko (1968) . For outgoing radiation

$$I = a + B T - (a_1 + B_1 T) n \quad (2)$$

where,

T : latitudinal surface temperature [°C]

n : fraction of cloudiness [-]

a : coefficient of value 14.0 [-]

B : coefficient of value 0.14 [-]

a₁ : coefficient of value 3.0 [-]

B₁ : coefficient of value 0.10 [-] .

For horizontal heat transfer in the atmosphere and hydrosphere

$$A = \beta (T - T_P) \quad (3)$$

where ,

T_P : mean earth temperature [°C]

β : coefficient of value 0.235 [kcal/cm² month °C] .

Equations (1), (2), and (3) are combined to form

$$T = \frac{Q(1 - \alpha) - G_E - a + a_1 + \beta T_P}{\beta + B - B_1 n} \quad (4)$$

T_P is found from

$$T_P = \frac{Q_p (1 - \alpha_p) - a + a_1 n}{B - B_1 n}$$

Q_p and α_p are average planetary values of radiation and albedo. In order to account for the variations of latitude belt surface area, the sampled latitude values of solar insolation and albedo were weighted with the cosine of the latitude when average planetary values were calculated.

The empirical Budyko equations, alone, cannot replicate the magnitude of historic surface temperature oscillations. Amplitude magnification of Budyko generated surface temperature was included in the model. For each sampled latitude, mean surface temperatures obtained from an exclusion of G_E simulation were read into the fortran

model from the data file AVGT.DAT. These mean surface temperatures were calculated from a one hundred thousand year interval. The differences between the model generated temperatures and these mean values were multiplied by an amplification factor and added to the mean values, $T_{GE_{amplified}} = (T_{GE} - AVGT) * amplification\ factor + AVGT$. The amplified surface temperatures were used as the forcing function in G_E calculations. Amplification for all simulations was a factor of ten, producing fluctuations on the order of ten to twenty degrees Celsius.

Solar Insolation

Variations of temporal and latitudinal solar radiation over the simulations were calculated using an algorithm from Berger (1979). The algorithm calculates the energy available at any given latitude, with the assumption of a perfectly transparent atmosphere. Berger also assumes a constant solar constant of 1353 W/m². Long term variations of the earth 's eccentricity, obliquity, and longitude of the perihelion are taken into account. The trigonometric series expansions

$$\varepsilon = \varepsilon^* + \sum_{i=1}^{i=47} A_i \cos(f_i t + \delta_i)$$

$$e \sin \omega = \sum_{i=1}^{i=47} P_i \sin(\alpha_i t + \zeta_i)$$

$$e \cos \omega = \sum_{i=1}^{i=47} P_i \cos(\alpha_i t + \zeta_i)$$

are utilized to find values for the eccentricity e , the perihelion ω , and the obliquity ε . In the expansions above, ε^* is 23.320556°, $t=0$ refers to 1950 A.D., t is negative for times prior to 1950 A.D. The sets of forty–seven coefficients to the right of the summation signs are read into the model code from data files. The files are A_i 's from OAMP.DAT, f_i 's from OMNRT.DAT, δ_i 's from OPHASE.DAT, P_i 's from EAMP.DAT, α_i 's from EMNRT.DAT,

and ζ_i 's from EPHASE.DAT. These files along with an independent code to calculate solar insolation given time and latitude, MILANK.FOR, are included in Appendix I.

Once the orbital parameters for a given time are known, the daily solar insolation for a given latitude and season can be calculated. The classical formulas for the daily insolation, W , are

1) for latitudes where there is no sunset

$$|\phi| \geq \frac{\pi}{2} - |\delta|$$

$$\phi > 0 \text{ if } \delta > 0$$

$$\phi < 0 \text{ if } \delta < 0$$

$$W = \frac{86.4S_0}{Q^2} \sin\phi \sin\delta$$

2) for latitudes where there is no sunrise

$$|\phi| \geq \frac{\pi}{2} - |\delta|$$

$$\phi < 0 \text{ if } \delta > 0$$

$$\phi > 0 \text{ if } \delta < 0$$

$$W = 0$$

3) for latitudes where there is daily sunset and sunrise

$$-\left(\frac{\pi}{2} - |\delta|\right) < \phi < \frac{\pi}{2} - |\delta|$$

$$W = \frac{86.4S_0}{\pi Q^2} (H_0 \sin\phi \sin\delta + \cos\phi \cos\delta \sin H_0)$$

H_0 is the absolute value of the hour angle at sunrise and sunset and is given by

$$\cos H_o = - \tan \phi \tan \delta \quad .$$

All the angles which locate the earth on its orbit are taken as being constant over the whole day. The declination is related to the true longitude of the sun by

$$\sin \delta = \sin \varepsilon \sin \lambda$$

The normalized earth 's sun distance is given by

$$e = \frac{r}{a} = \frac{1 - e^2}{1 + e \cos \nu} \quad .$$

Definitions for parameters in the above equations are

ϕ : latitude,

δ : declination of the sun,

ρ : earth–sun distance r measured in units of the semimajor axis a ,

H : hour angle of the sun during the day,

ν : true anomaly: positional angle of the earth on its orbit, counted counter clockwise from the perihelion,

M : mean anomaly: positional angle of a "mean" earth rotating around the sun with a constant angular speed equal to $2\pi/T$ and counted counterclockwise from the perihelion,

λ : true longitude of the earth is counted counterclockwise from the vernal equinox and is related to ν through $\lambda = \nu + \omega$. As in this formula, ω is measured from the vernal equinox,

λ_m : mean longitude associated with the mean earth is related to

M : $\lambda_m = M + \omega$.

In the model, latitudes were sampled from 0° to 90° with 15° intervals.

For each sample latitude, the solar insolation values were calculated at twenty–four equal

subdivisions of each year. The solar insolation values were then averaged over the year, for each sample latitude, and used in the energy balance calculations..

Albedo

Albedos, for each latitudinal increment, were chosen based on present planetary climate conditions. For latitudes between 0° and 60° an average albedo of 0.32 was chosen. For latitudes between 60° and 70° the albedo was increased linearly from 0.32 to 0.50. For latitudes from 70° to 90° the albedo was increased linearly from 0.50 to 0.74. Constant albedo values, independent of surface temperature, were used to lend higher stability to the model. Energy balance models with transient albedos often cover the earth with an expanded icecap after a few thousand years (Schneider, Gal–Chen, 1973). A transient albedo function may be accounted for in future adjustments to the model.

Change in solid earth heat storage

The change in solid earth heat storage, G_E , if written in integral form is

$$G_E = C \int_0^d \frac{\partial T_E}{\partial t} dz \quad (5)$$

where,

C : Heat capacity of solid earth [kcal/cm³ °C]

d : depth in solid earth of lower boundary condition [cm]

T_E : Temperature at depth , z , for time , t , in subsurface earth [°C]

t : time [months]

z : depth [cm] .

Subsurface temperature distributions are determined using the heat flow continuity equation. For an isotropic, homogeneous body without interior heat sources or sinks, the one dimensional partial differential equation for transient heat conduction in the subsurface earth is

$$D_T \frac{\partial^2 T_E}{\partial z^2} = \frac{\partial T_E}{\partial t} \quad (6)$$

where, D_T is the thermal diffusivity constant [cm^2/month]. In the model, this equation is evaluated numerically using an implicit finite difference approach. This approach generates the system of equations

$$\begin{bmatrix} 1+\lambda & -\lambda/2 & 0 & 0 & \dots & \dots & \dots & \dots & 0 \\ -\lambda/2 & 1+\lambda & -\lambda/2 & 0 & \dots & \dots & \dots & \dots & 0 \\ 0 & -\lambda/2 & 1+\lambda & -\lambda/2 & 0 & \dots & \dots & \dots & 0 \\ 0 & 0 & -\lambda/2 & 1+\lambda & -\lambda/2 & 0 & \dots & \dots & 0 \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 0 & 0 & \dots & \dots & \dots & 0 & -1 & 1 & \dots \end{bmatrix} \begin{bmatrix} T_{E_1}^{j+1} \\ T_{E_2}^{j+1} \\ T_{E_3}^{j+1} \\ \vdots \\ \vdots \\ \vdots \\ \vdots \\ T_{E_m}^{j+1} \end{bmatrix} = \begin{bmatrix} (1-\lambda)T_{E_1}^j + \lambda/2 (T_{E_2}^j + T_{E_0}^j + T_{E_0}^{j+1}) \\ (1-\lambda)T_{E_2}^j + \lambda/2 (T_{E_1}^j + T_{E_3}^j) \\ (1-\lambda)T_{E_3}^j + \lambda/2 (T_{E_2}^j + T_{E_4}^j) \\ \vdots \\ \vdots \\ (1-\lambda)T_{E_{m-1}}^j + \lambda/2 (T_{E_{m-2}}^j + T_{E_m}^j) \\ (.03 \text{ }^\circ\text{C/m})\Delta z \end{bmatrix} \quad (7)$$

Above, $\lambda = D_T \Delta t / (\Delta z)^2$, the superscript $j+1$ is the present time step, and the superscript j is the previous time step. The subscript on T_E is the depth node position. T_{E_0} represents surface temperature. Boundary conditions were included in the system of equations. For each

time step, the boundary condition at the surface of the earth was determined from equation (4). The G_E determined from the previous time iteration was used in the equation (4) calculation. The lower boundary condition was the geothermal gradient. A thermal gradient of .03 °C/meter was used (Lapidus).

The system of equations were solved using a Gauss–Seidal iterative algorithm. Whenever possible, this algorithm makes use of newly computed values of T_{Ei} to obtain greater efficiency in convergence to the solution. The calculation

$$T_{E_i} = \frac{- \sum_{l=1}^{l=i-1} a_{il} T_{E_l} - \sum_{l=i+1}^{l=m} a_{il} PT_{E_l} + b_i}{a_{ii}}$$

is preformed until $T_{Ei} - PT_{Ei}$ is below a prescribed tolerance level. PT_{Ei} is depth temperature at node i from the previous iteration. The tolerance level used in the model was 10^{-6} °C. The terms a_{il} and b_i are matrix positions in the matrixes on the left and right sides of equation (7), respectively. At the start of each Gauss–Siedal subroutine call, the initial values of PT_{Ei} , were depth node temperatures from the previous time step.

Numerical Differentiation

For equation (5), the derivatives of subsurface temperature with respect to time for a given depth were evaluated using a three point formula with Lagrange coefficient polynomials. Subroutine SUBSURF uses the equation

$$\left(\frac{dT}{dt} \right)_{at \ t} = - \frac{1}{\Delta t} \left[- \frac{3}{2} T_{at \ t} + 2T_{at \ t-\Delta t} - \frac{1}{2} T_{at \ t-2\Delta t} \right]$$

to calculate these derivatives. The three point formula lent greater accuracy to the model than would a two point formula.

Numerical Integration

The integration of these derivatives with respect to depth was accomplished using a Simpson Composite numerical integration scheme. This method subdivides the interval [a, b] into n subintervals and uses Simpson 's rule on each pair of consecutive subintervals, leading to improved precision. Subroutine SUBSURF uses the equation

$$\int_a^b f(x)dx = \sum_{j=1}^{j=m} \frac{4x}{3} [f(x_{2j-2}) + 4f(x_{2j-1}) + f(x_{2j})]$$

to accomplish the Simpson Composite numerical integration. In this equation, $m=n/2$, where n is the number of depth nodes.

All calculations in the model were done in double precision. The fortran code representing the model is in Appendix II.

Model Verification

To check the model, a heat transfer boundary value problem with known analytic solution was utilized. The problem that was tested was

$$D_T \frac{\partial^2 T_E}{\partial z^2} = \frac{\partial T_E}{\partial t}, \quad 0 < z < d, \quad 0 < t$$

$$T_E(0,t) = T_0, \quad 0 < t$$

$$T_E(d,t) = T_1, \quad 0 < t$$

$$T_E(z,0) = T_1, \quad 0 < z < d.$$

The analytic solution is $T_E(z,t) = T_0 + (T_1 - T_0) \frac{z}{d} + \sum_{n=1}^{n=\infty} b_n \sin \lambda_n z e^{-\lambda_n^2 D_T t}$

where, $b_n = \frac{2}{n\pi} (T_1 - T_0)$ and $\lambda_n = \frac{n\pi}{d}$. After differentiating the analytic solution for T_E with respect to t and integrating with respect to z, the analytic solution for change in subsurface heat storage was obtained,

$$G_E = C \int_0^d \frac{\partial T_E}{\partial t} dz = C \sum_{n=0}^{n=\infty} \lambda_n D_T b_n (\cos \lambda_n d - 1) e^{-\lambda_n^2 D_T t} .$$

A comparison with the numerical model output yielded identical solutions.

Results

The numerical model was used to simulate the effect of inclusion of earth heat storage into the earth surface energy balance calculation. Omission of earth heat storage in the energy balance model is equivalent to an earth thermal diffusivity of zero. Deviations in surface temperature from this $D_T=0$ case were observed over a range of plausible solid earth thermal diffusivities. A range of reasonable values for heat capacity and thermal diffusivity of the earth were obtained from Saltzman(1985), Oliver and Fairbridge(1987), and James(1989). Values for some common earth materials are shown in table 1.

table 1. Thermal parameters of various earth materials
(after Oliver and Fairbridge)

material	heat capacity cal/cm ³ °C	thermal diffusivity cm ² /sec
ice	0.45	0.012
dry sand	0.3	0.0013
wet soil	0.4	0.01
still water	1.0	0.0015
stirred water	1.0	50 variable and estimated

Simulations were performed for eight values of thermal diffusivity. These values ranged from 2×10^{-6} m²/sec, a best estimate of solid earth thermal diffusivity, to 200×10^{-6} m²/sec. For diffusivity values in this range, a depth node spacing of approximately one hundred meters produced results nearly identical to depth node spacings of either ten

meters or one meter. The greater node spacing was chosen for its rapidity of calculation. Time node spacing of fifty years yielded stable output.

Figure 2 demonstrates the lag and magnitude of the change in heat storage oscillations for the total depth $d=1000$ meters case. For comparison purposes, the $D_T=0$ average planetary temperature case is illustrated in the upper plot. G_E is not an exact linear transformation of ΔT_P . Visually, however, the differences between the curves are indistinguishable when a second y-axis is added to the lower plot. The magnitude of these G_E peaks translates to a maximum temperature difference of about five hundredth of a degree Centigrade when compared to the exclusion of heat storage, $D_T=0$, case. As shown in figure 3, the peaks asymptotically approach a limiting value as the thermal diffusivity increases. This limiting value is dependant on the depth to the lower boundary condition, d . For $d=3000$ meters, the energy ratio approaches a much larger limiting value. This is also shown in figure 3. At the lower end of the diffusivity range, the differences between the energy ratios are minor.

Figure 4 and table 3 reveal the lag structure of G_E with variations in thermal diffusivity and depth, d . Again, for D_T 's near the best estimate end of the axis, there is little variation with changes in effective penetration depth. Differences in the lag structure and energy ratio curves may be due to energy balance contributions from the constant energy flux at the lower boundary condition.

The contribution of changes in solid earth heat storage into the surface energy balance appear to be insignificant when thermal diffusivity values near $2 \times 10^{-6} \text{ m}^2/\text{sec}$ are taken as average earth values. To determine the effects on surface temperature oscillations if changes in solid earth heat storage were a significant factor, the model was executed using a thermal diffusivity value of $.005 \text{ m}^2/\text{second}$. This value is the estimated thermal diffusivity of stirred water. In order to maintain stable calculations, depth node spacing of 3000 meters was required. The total depth was 27,000 meters. The time node spacing was 50 years. Figures 5 and 6 reveal the output of the model calculation. The time

range was set for a period of 1.2 million years for comparison with figure 1. The output signal with inclusion of G_E is a dampening of the noninclusive signal. Lags between the two temperature outputs are less than 2 kyr. E-fold response time for a step input is given by $t=d^2/D_T$ (James, 1989) . A value of 4.62 kyr is obtained using the above parameters.

Comparison spectral plots are in figure 7 and figure 8. Figure 7 is created from spectral data from the exclusion of the G_E term case. Figure 8 is created from spectral data from the inclusion of G_E case. The contour plot data files were formed using moving 200,000 year subintervals spaced 40,000 years apart. The spectral calculations were performed using a Lomb periodogram. The fortran code which created the data files is PLATE.F. It is included in appendix III.

Spectral contour plots reveal only minor contrasts between the two cases. There is a slight elongation in the spectral mound near frequency .052 cycles/kyr (period of 19,200 years) and the time 400 to 500 ka. There is also some change in the low power mounds near frequency .01 cycles/kyr (period of 100,000 years). Inclusion of G_E into the energy balance facilitates the loss of the mound near the time 520 to 690 ka. The inclusion of G_E also reshapes the mound from 930 to 1080ka. Overall however, it appears that as the uniform thermal diffusivity chosen for the earth is increased and the lag between G_E and Q_P goes to zero, the two temperature cases become nearly linearly related.

Also of note, in figures 7 and 8, are the absence of spectral mounds corresponding to orbital forcing due to variations of the obliquity of the ecliptic. Orbital forcing of this variety should generate frequencies near .0243 cycles/kyr (periods of 41,000 years). The averaging process, which weights lower latitude solar insolation values more heavily, was thought to dampen out these frequencies. Figures 9 and 10, spectral contour plots of surface temperature values simulated at a latitude of 75° , demonstrate activity in the obliquity frequency area. As in the average surface temperature spectral plots, inclusion of G_E into the energy balance equation does not significantly alter the frequency pattern of the exclusion of G_E case.

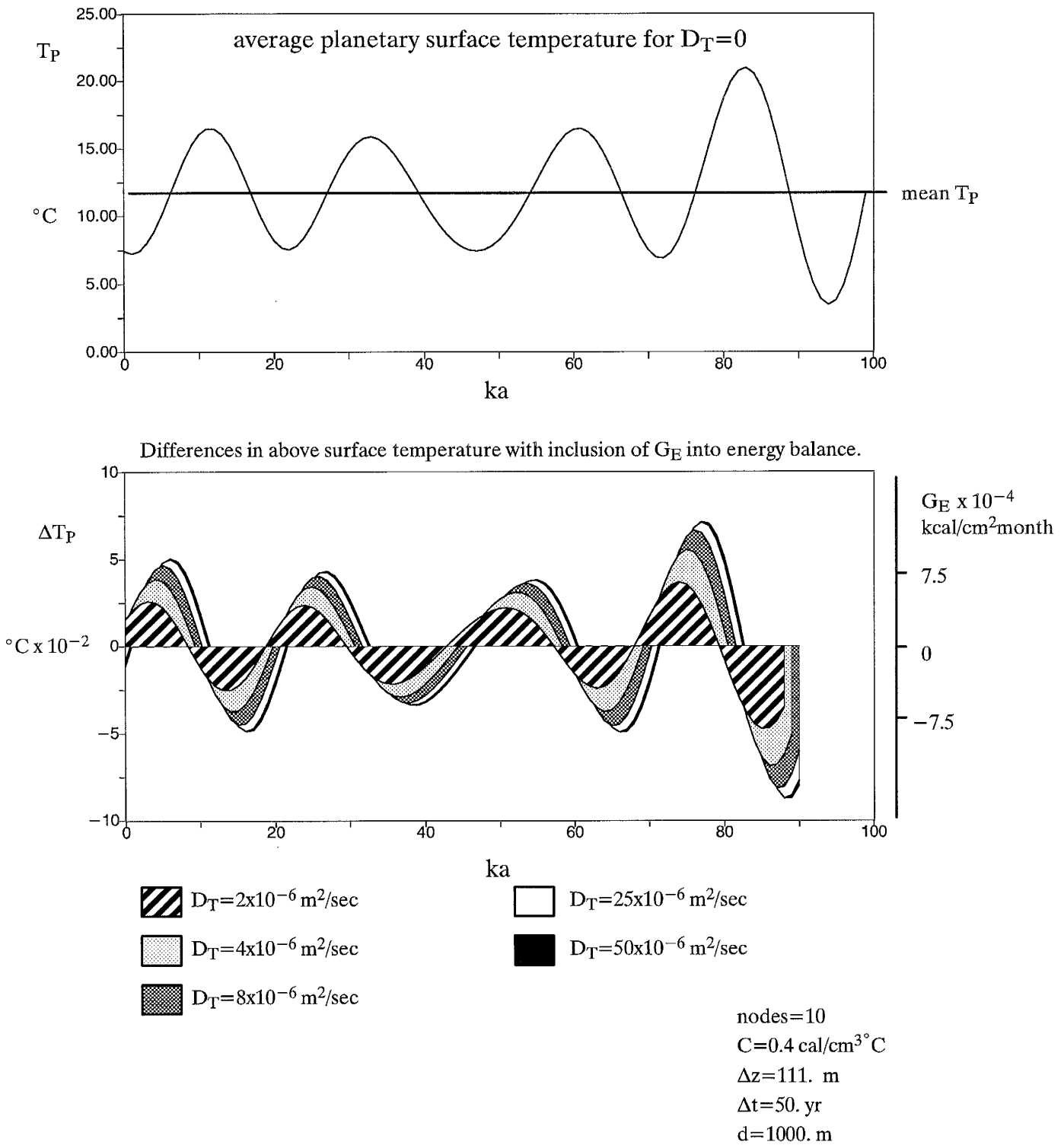


Figure 2. Effect of varying earth thermal diffusivities on subsurface energy contribution.

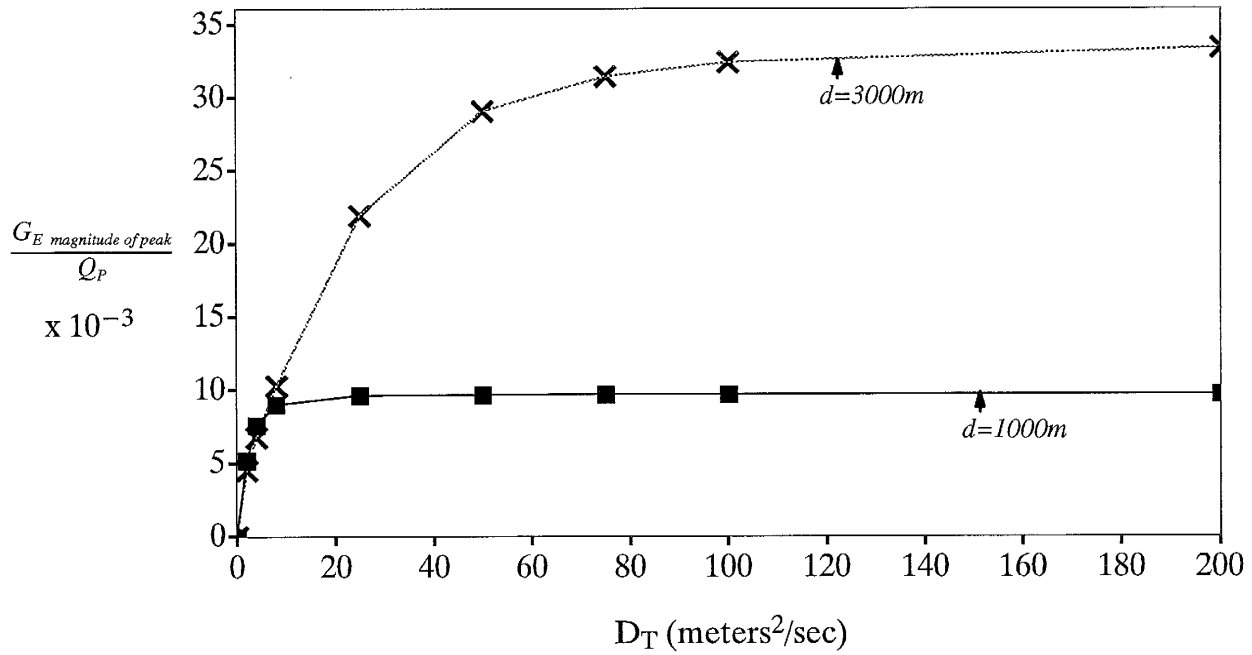


Figure 3. Sensitivity plot of change in energy ratio with variation of thermal diffusivity.

Table 2. variation of average ratio with thermal diffusivity.

D_T (m ² /sec)	d=1000 m	d=3000 m
	ratio x 10 ⁻³	ratio x 10 ⁻³
2 x 10 ⁻⁶	5.199	4.518
4 x 10 ⁻⁶	7.604	6.775
8 x 10 ⁻⁶	9.015	10.276
25 x 10 ⁻⁶	9.6125	21.888
50 x 10 ⁻⁶	9.666	29.066
75 x 10 ⁻⁶	9.675	31.441
100 x 10 ⁻⁶	9.677	32.406
200 x 10 ⁻⁶	9.681	33.381

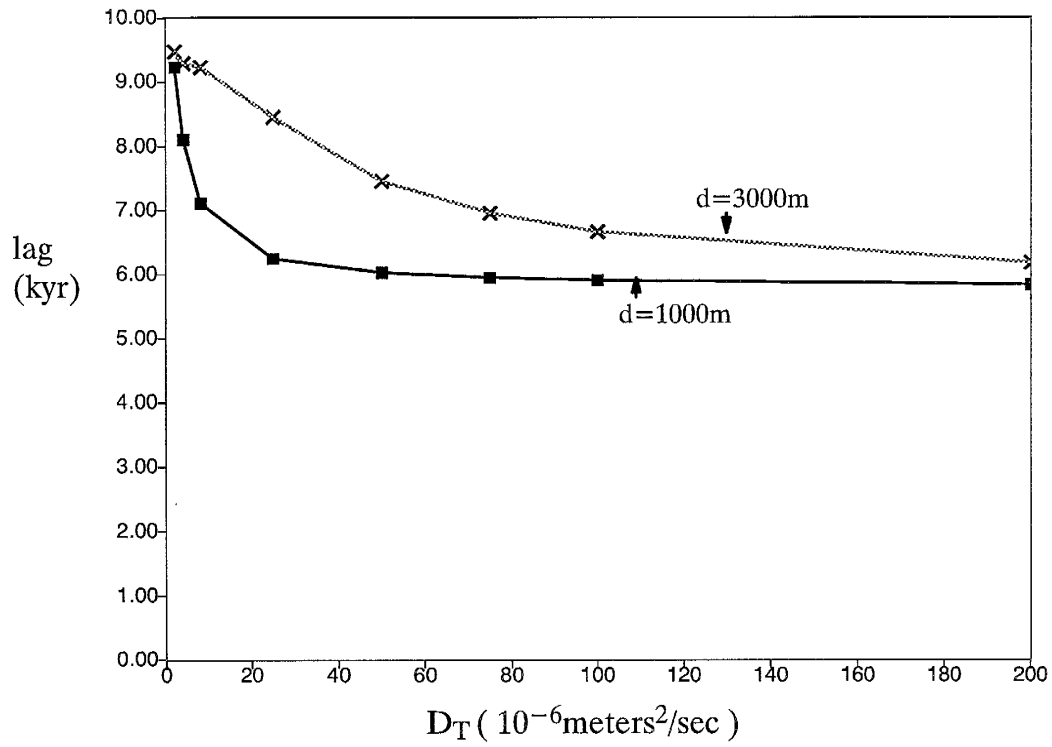


Figure 4. Sensitivity of lag with changes in thermal diffusivity.

Table 3. variation of average lag with thermal diffusivity.

D_T (m^2/sec)	$d=1000 \text{ m}$	$d=3000 \text{ m}$
	avg. lag (kyr)	avg. lag (kyr)
2×10^{-6}	9.225	9.47
4×10^{-6}	8.100	9.29
8×10^{-6}	7.100	9.225
25×10^{-6}	6.240	8.45
50×10^{-6}	6.025	7.45
75×10^{-6}	5.950	6.95
100×10^{-6}	5.906	6.66
200×10^{-6}	5.837	6.18

figure 5. Comparison of average surface temperature over 1200 kyr.

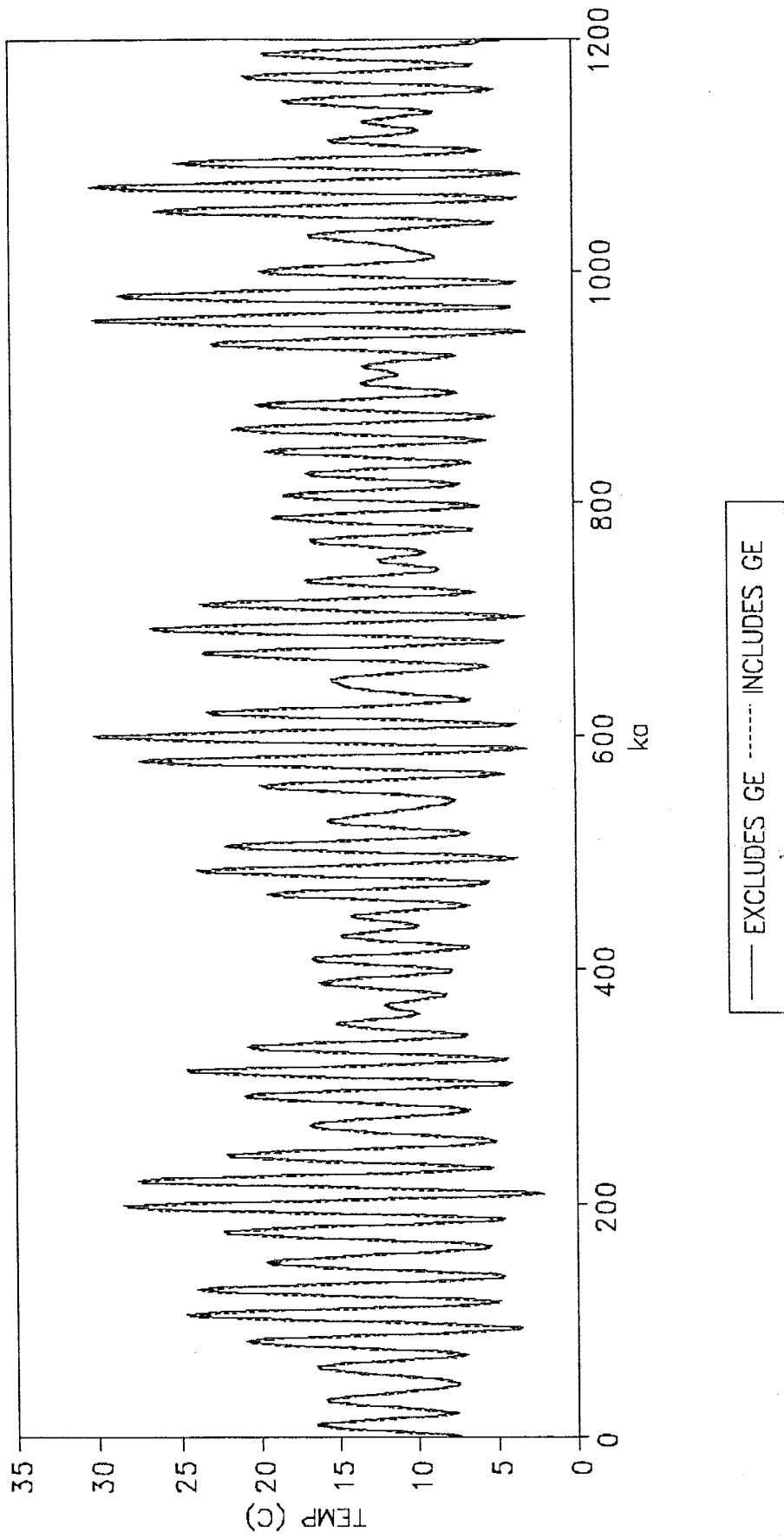


figure 6. Comparison of average surface temperature over 200 kyr.

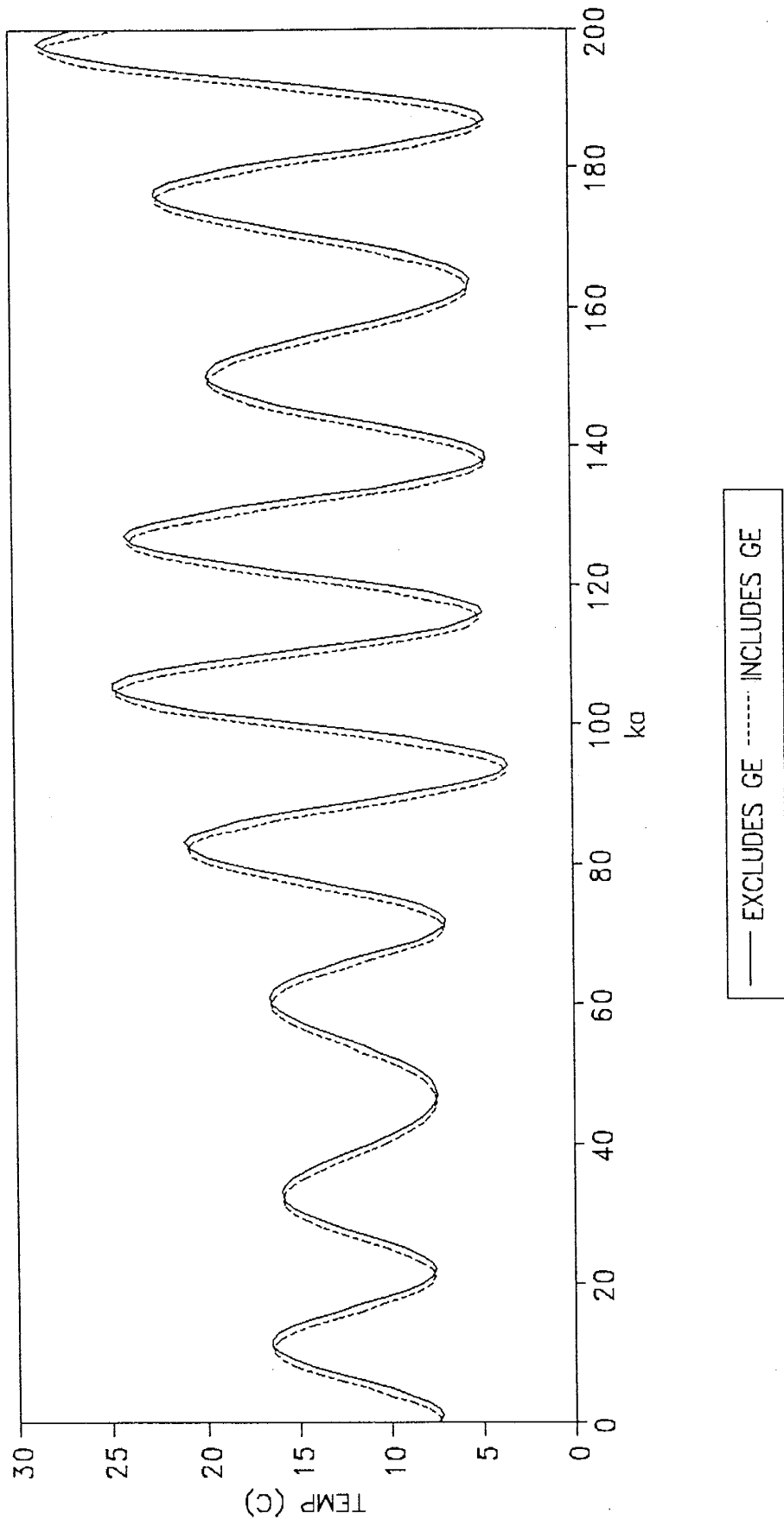


figure 7. Spectral density contour plot of average earth surface temperature, excluding G_E .

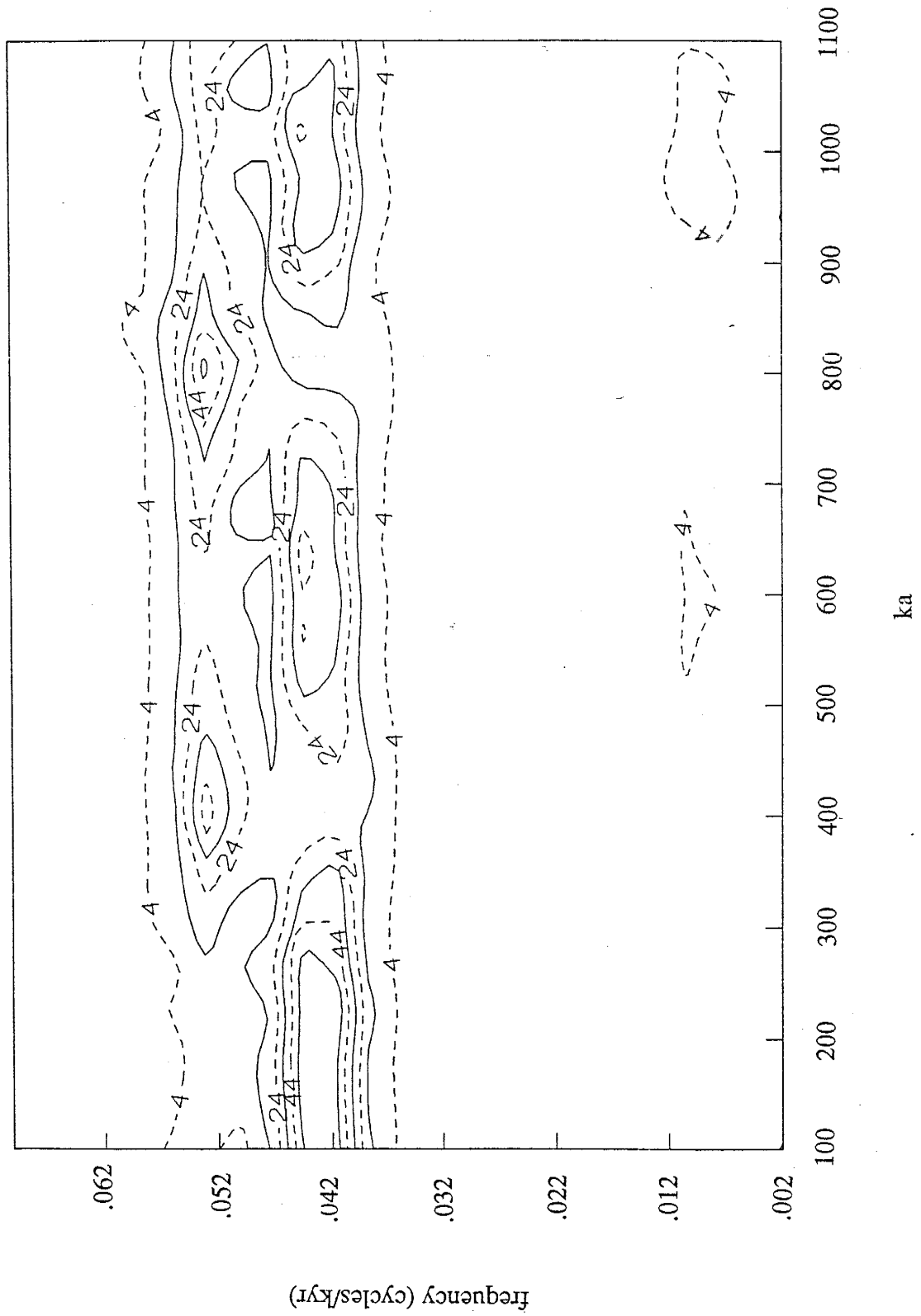
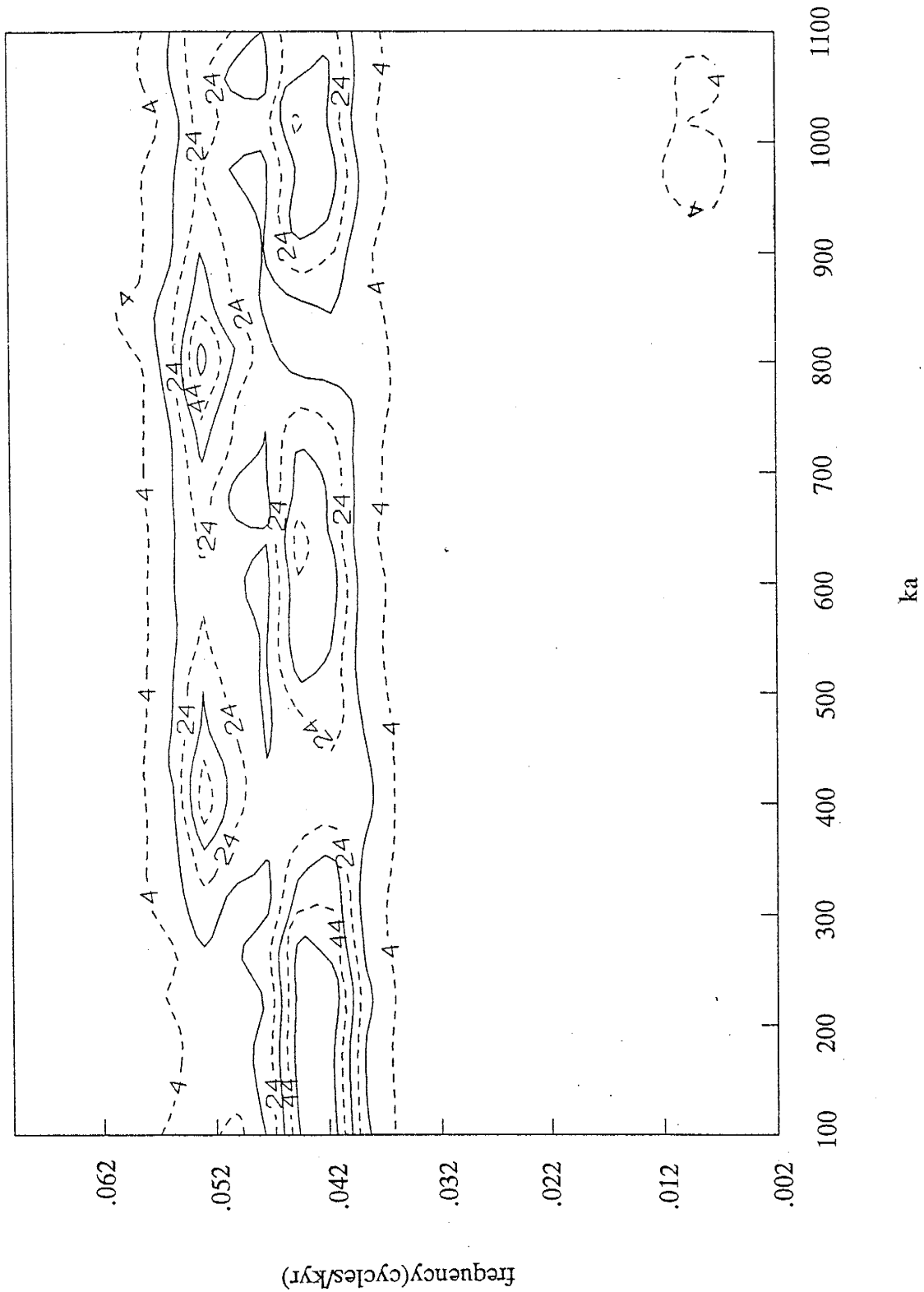


figure 8. Spectral density contour plot of average earth surface temperature, including G_E .



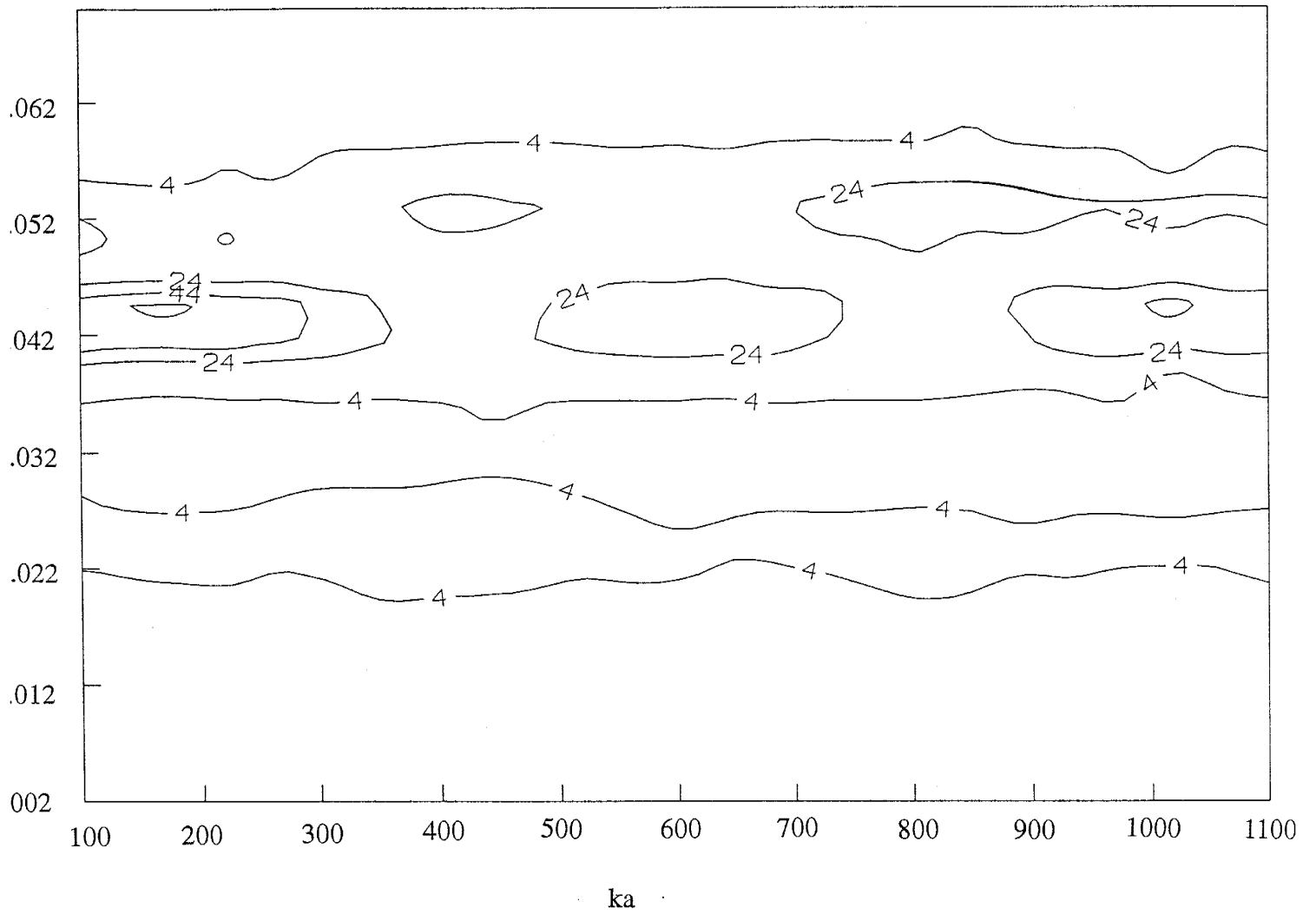


figure 9. Spectral density contour plot of earth surface temperature at 75° latitude, excluding G_E.

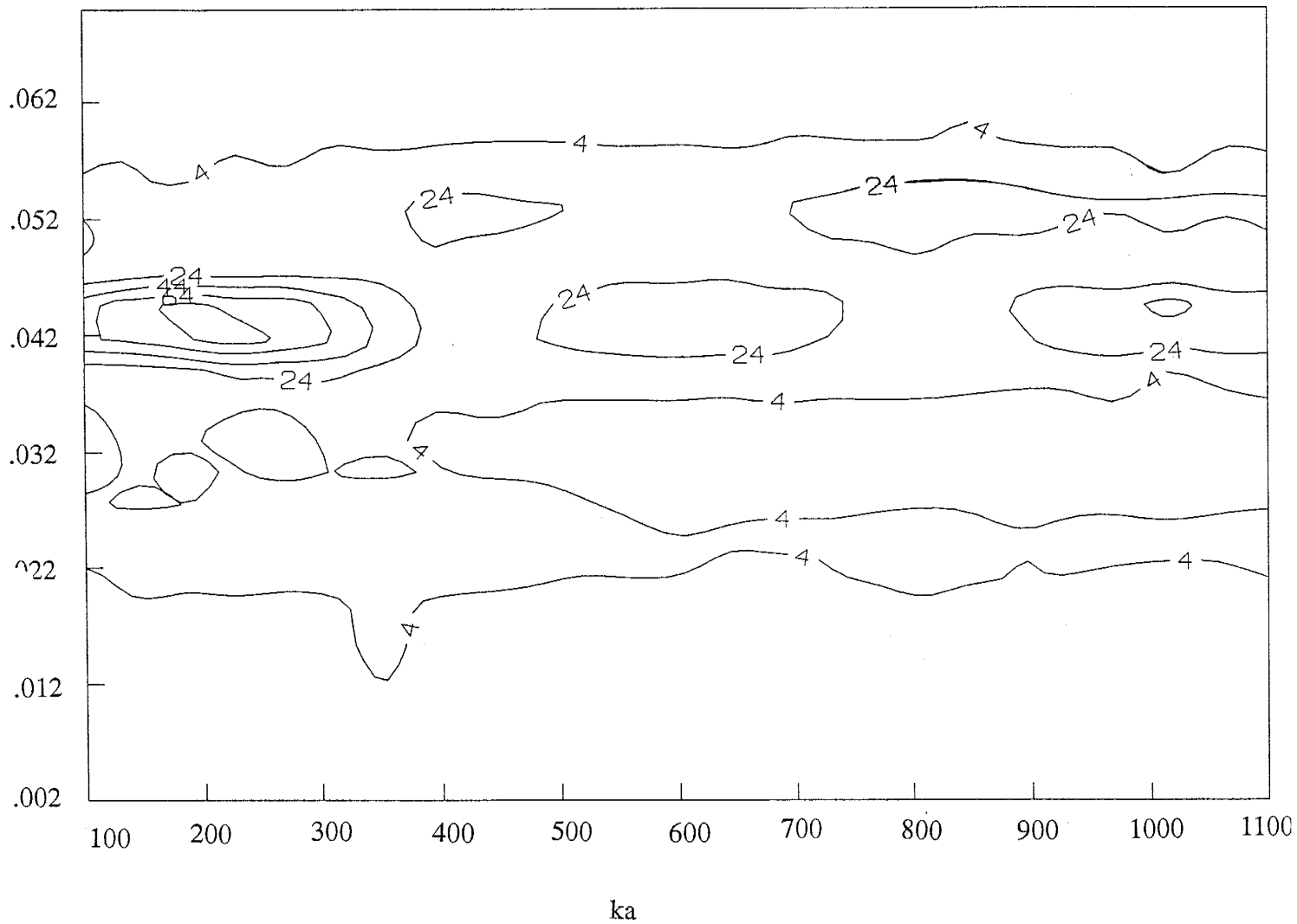


figure 10. Spectral density contour plot of earth surface temperature at 75° latitude, including G_E .

Conclusions

The model demonstrates that changes in solid earth heat storage represent a very small portion of the total energy budget at the surface of the earth. The choice of effective penetration depth effects maximum energy ratio. For values of thermal diffusivity on the order of 10^{-6} m²/sec, however, the solid earth heat storage contribution represents a Budyko temperature deviation of only .05°C.

Forcing the thermal diffusivity to magnitudes large enough to produce significant temperature deviations, energy contributions from changes in earth heat storage no longer lag far enough behind the solar insolation contribution to produce nonlinearity in the system. More complex thermal couplings would undoubtable occur if separate land and ocean masses were discretized in the model. Future versions of the model will also require simulation of ice cap expansion, and the consequent effect on albedo.

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APPENDIX I

```
*****
CALCULATES SOLAR INSOLATION, USING AN ALGORITHM FROM BERGER, A. L.,*
'LONG-TERM VARIATIONS OF DAILY INSOLATION AND QUATERNARY CLIMATE '*
CHANGES,' JOURNAL OF THE ATMOSPHERIC SCIENCES, 35:2362-2367. *
*****
```

```
DOUBLE PRECISION sumob, pi, sumc1, sumc2, obliq, obconst,
+ long, eccen, oamp(50), omnrt(50), ophase(50),
+ eamp(50), emnrt(50), ephase(50), const1, const2, solrconst,
+ phi, delta, sq, one, solrins,rho,insconst,
+ Ho,two,sqrt,LAMD,lam,nu
```

```
parameter (obconst= 23.320556, pi = 3.141592653, const1 = 3600.,
+ const2 = 180., solrconst=1.95, insconst=1440., sq=2.,
+ one=1., two = 2., sqrt = 0.5)
```

```
** All of the solar insolation values a calculated for the month **
** corresponding to the parameter 'lam' . Lam=0 corresponds **
** to approx March 21. The mid month values are defined by a **
** change in 'lam' of thirty(30) degrees, i.e., april=30,may=60. **
```

```
** read coeff. into arrays **
** These coeff. are used to calculate the earths' wobbles and tilts, **
** or more specifically, its eccentricity, obliquity, and longitude **
** of the perihelion. **
```

```
open(unit=61, status = 'old', file = 'oamp.dat')
read (61,*) (oamp(i), i = 1,47)
write(6,*)' read ',i-1,' points from oamp'
close(unit=61)
```

```
open(unit=62, status = 'old',file = 'omnrt.dat')
read (62,*)(omnrt(j), j = 1,47)
write(6,*)' read ',j-1,' points from omnrt'
close(unit=62)
```

```
open(unit=63, status = 'old',file = 'ophase.dat')
read (63,*)(ophase(k), k = 1,47)
write(6,*)' read ',k-1,' points from ophase'
close(unit=63)
```

```
open(unit=64, status = 'old',file = 'eamp.dat')
read (64,*)(eamp(i), i = 1,46)
write(6,*)' read ',i-1,' points from eamp'
close(unit=64)
```

```
open(unit=65, status = 'old',file = 'emnrt.dat')
read (65,*)(emnrt(j), j = 1,46)
write(6,*)' read ',j-1,' points from emnrt'
close(unit=65)
```

```
open(unit=66, status = 'old',file = 'ephas.dat')
read (66,*)(ephas(k), k = 1,46)
write(6,*)' read ',k-1,' points from ephase'
close(unit=66)
```

```
WRITE(6,*)'ALL INPUT DATA READ'
```

```

WRITE(6,*)'ENTER VALUE FOR STARTING TIME'
WRITE(6,*)'-10000. MEANS TEN THOUSAND YEARS AGO '
WRITE(6,*)
READ(5,*)TIME
WRITE(6,*)

```

```

WRITE(6,*)'ENTER VALUE FOR LATITUDE IN DEGREES'
WRITE(6,*)
READ(5,*)PHID
WRITE(6,*)

```

```

WRITE(6,*)'ENTER VALUE FOR SEASON'
WRITE(6,*)
WRITE(6,*)'MARCH 21=0., JUNE 21=90., SEPTEMBER 21=180., ETC.'
WRITE(6,*)
READ(5,*)LAMD
WRITE(6,*)

```

```

** Convert to radians **

```

```

phi = phiD*(pi/const2)
lam = lamD*(pi/const2)

```

```

C.....

```

```

sumc1 = 0.
sumc2 = 0.
sumob = 0.

```

```

do 200 j = 1,47

```

```

** Mid-month obliquity for a given LATITUDE AND TIME OF YEAR **

```

```

+ sumob = sumob + (oamp(j)/const1)*(cos(((omnrt(j)*dble(time))
+ /const1 + ophase(j))*(pi/const2)))

```

```

** sumc1 & sumc2 are used to calculate eccentricity and longitude of the **
** perihelion based on a moving vernal equinox **

```

```

+ sumc1 = sumc1 + eamp(j)*sin(((emnrt(j)*dble(time))/const1
+ + ephase(j))*(pi/const2))

```

```

+ sumc2 = sumc2 + eamp(j)*cos(((emnrt(j)*dble(time))/const1
+ + ephase(j))*(pi/const2))

```

```

200 continue

```

```

obliq = sumob + obconst
long = (atan2(sumc1,sumc2))*(const2/pi)
eccen = sumc1/sin(atan2(sumc1,sumc2))

```

```

** PARAMETERS USED TO CALCULATE MID-MONTH SOLAR INSOLATION **

```

```

delta = asin((sin(lam))*sin(obliq*(pi/const2)))

```

```

nu = lam-((long+const2)*(pi/const2))

```

```

+ rho = (one-(eccen**sq))/
+ (one+(eccen*cos(nu)))

```



```
** MID-MONTH SOLAR INSOLATION IN CAL/CM2-DAY **
```

```
IF(ABS(PHI).GE.(PI/TWO-abs(DELTA)))THEN  
+   if(((phi.gt.(0.)).and.(delta.gt.(0.))).or.((phi.lt.(0.)).  
+     and.(delta.lt.(0.))))then  
+     SOLRINS=      ((insconst*solrconst)/(rho**sq))  
+       *sin(phi)*sin(delta)  
+     end if  
+   if(((phi.lt.(0.)).and.(delta.gt.(0.))).or.((phi.gt.(0.)).  
+     and.(delta.lt.(0.))))then  
+     SOLRINS = DBLE(0.)  
+     end if  
+   ELSE  
+     Ho = acos(-tan(phi)*tan(delta))  
+     solrins = ((insconst*solrconst)/(pi*(rho)**sq))  
+       *(Ho*sin(phi)*sin(delta)+cos(phi)*cos(delta)*sin(Ho))  
+     END IF  
+   WRITE(6,*)'SOLAR INSOLATION = ',SOLRINS,' CAL/CM2 DAY'  
+   END
```

EAMP.DAT

0.018608
0.0162752
-0.0130066
0.0098883
-0.003367
0.0033308
-0.002354
0.0014002
0.001007
0.000857
0.0006499
0.000599
0.000378
-0.000337
0.0003334
0.0003334
0.0002916
0.0002916
0.000276
-0.000233
-0.000233
0.000182
0.0001772
0.0001772
-0.0001740
-0.000124
0.0001153
0.0001153
0.0001008
0.0001008
0.0000912
0.0000912
-0.0000806
-0.0000806
0.0000798
0.0000798
-0.0000638
-0.0000638
0.0000612
0.0000612
-0.0000603
-0.0000603
0.0000597
0.0000597
0.0000559
0.0000559

EMNRT.DAT

54.646484
57.785370
68.296539
67.659821
67.286011
55.638351
68.670349
76.656036
56.798447
66.649292
53.504456
67.023102
68.933258
56.630219
86.256454
23.036499
89.395340
26.175385
69.307068
99.906509
36.686569
67.864838
99.269791
36.049850
56.625275
68.856720
87.266983
22.025970
90.405869
25.164856
78.818680
30.474274
100.917038
35.676025
81.957565
33.613159
92.468735
44.124329
100.280319
35.039322
98.895981
35.676025
87.248322
24.028381
86.630264
22.662689

EPHASE.DAT

32.01
197.18
311.69
323.59
282.76
90.58
352.52
131.83
157.53
294.66
118.25
335.48
299.80
149.16
283.91
320.11
89.08
125.27
340.62
203.60
239.79
155.48
215.49
251.68
130.23
214.05
312.84
291.17
118.01
96.34
160.31
83.70
232.53
210.86
325.48
248.87
80.00
3.39
244.42
222.75
174.67
210.86
342.48
18.68
324.73
279.28

OAMP.DAT

-2462.22
 -857.32
 -629.32
 -414.28
 -311.76
 308.94
 -162.55
 -116.11
 101.12
 -67.69
 24.91
 22.58
 -21.16
 -15.65
 15.39
 14.67
 -11.73
 10.27
 6.49
 5.85
 -5.49
 -5.43
 5.16
 5.08
 -4.07
 3.72
 3.40
 -2.83
 -2.66
 -2.57
 -2.47
 2.46
 2.25
 -2.08
 -1.97
 -1.88
 -1.85
 1.82
 1.76
 -1.54
 1.47
 -1.46
 1.42
 -1.18
 1.18
 -1.13
 1.09

OMNRT.DAT

31.60997
 32.620499
 24.172195
 31.983780
 44.828339
 30.973251
 43.668243
 32.246689
 30.599442
 42.68132
 43.836456
 47.439438
 63.219955
 64.230484
 1.01053
 7.437771
 55.782181
 .373813
 13.218362
 62.583237
 63.593765
 76.438309
 45.815262
 8.448301
 56.792709
 49.747849
 12.058272
 75.278214
 65.241013
 64.604294
 1.647247
 7.811584
 12.207832
 63.856659
 56.155991
 77.448837
 6.801054
 62.209412
 20.656128
 48.344406
 55.145462
 69.000534
 11.071350
 74.291306
 11.047742
 .636717
 12.844549

OPHASE.DAT

251.9
 280.83
 128.3
 292.72
 15.37
 263.79
 308.42
 240.
 222.97
 268.78
 316.79
 319.6
 143.8
 172.73
 28.93
 123.59
 20.2
 40.82
 123.47
 155.69
 184.62
 267.27
 55.01
 152.52
 49.13
 204.66
 56.52
 200.32
 201.66
 213.55
 17.03
 164.41
 94.54
 131.91
 61.03
 296.2
 135.48
 114.87
 247.06
 256.61
 32.1
 143.68
 16.87
 160.68
 27.59
 348.1
 82.64

APPENDIX II

```

*****
* GENERATES TRANSIENT SOLAR INSOLATION VALUES AT EARTH'S EQUATOR . *
* USES THESE SOLAR INSOLATION VALUES TO DRIVE A SIMPLE BUDYKO *
* MODEL WHICH DETERMINES TEMPERATURE VALUES AT A GIVEN LATITUDE. ALBEDO *
* VALUES ARE INVARIANT OF TEMPERATURE . FEEDBACK FROM SOLID EARTH *
* HEAT STORAGE IN THE ENERGY BALANCE IS ACCOUNTED FOR VIA A *
* SUBSURFACE IMPLICIT FINITE DIFFERENCE APPROXIMATION . *
*****

```

```

DOUBLE PRECISION sumob, pi, sumc1, sumc2, obliq, obconst,
+ long, eccen, oamp(50), omnrt(50), ophase(50), ESINW50,
+ eamp(50), emnrt(50), ephase(50), const1, const2, solrconst,
+ phi, delta, sq, one, solrins, rho, GRAD(7,1000), insconst,
+ Ho, two, trop, sqrt, TS(7), Q(7), ALBEDO(7), LAMD, lam, C, nu,
+ DELTAZ, TEMP(7,1000), TLAST(7,1000), T2AGO(7,1000), DT,
+ LAMDINT, OBL50, BETA, A1, A2, B1, B2, GE(7),
+ TP, QP, SUMLATQ, SUMCOS, SUMGE, AVG_GE, LASTGE(7),
+ TNOGE(7), TPGE, AMP, TAMP(7), AVGT(7), TSAMP(7), SUMTAMP,
+ SUMTSAMP, AVG_TAMP, AVG_TSAMP, deltat, weight, sumweight

```

INTEGER NODES

```

parameter (obconst= 23.320556, pi = 3.141592653, const1 = 3600.,
+ const2 = 180., solrconst=1.95, insconst=1440., sq=2.,
+ one=1., two = 2., sqrt = 0.5,
+ trop = 525948.761666, obl50=23.44631, BETA=.235,
+ esinw50 = .016109170339, A1=14.0, A2=3.0, B1=.14, B2=.10)

```

```

** All of the solar insolation values a calculated for the month **
** corresponding to the parameter 'lam' . Lam=0 corresponds **
** to approx March 21. The mid month values are defined by a **
** change in 'lam' of thirty(30) degrees, i.e., april=30, may=60. **

```

```

** read coeff. into arrays **
** These coeff. are used to calculate the earths' wobbles and tilts, **
** or more specifically, its eccentricity, obliquity, and longitude **
** of the perihelion. **

```

```

open(unit=60, status = 'old', file = 'avgt.dat')
read (60,*) (AVGT(i), i = 1,7)
write(6,*)' read ',i-1,' points from AVGT'
close(unit=60)

```

```

open(unit=61, status = 'old', file = 'oamp.dat')
read (61,*) (oamp(i), i = 1,47)
write(6,*)' read ',i-1,' points from oamp'
close(unit=61)

```

```

open(unit=62, status = 'old', file = 'omnrt.dat')
read (62,*)(omnrt(j), j = 1,47)
write(6,*)' read ',j-1,' points from omnrt'
close(unit=62)

```

```

open(unit=63, status = 'old', file = 'ophase.dat')
read (63,*)(ophase(k), k = 1,47)
write(6,*)' read ',k-1,' points from ophase'
close(unit=63)

```

```

open(unit=64, status = 'old',file = 'eamp.dat')
read (64,*)(eamp(i), i = 1,46)
write(6,*)' read ',i-1,' points from eamp'
close(unit=64)

open(unit=65, status = 'old',file = 'emnrt.dat')
read (65,*)(emnrt(j), j = 1,46)
write(6,*)' read ',j-1,' points from emnrt'
close(unit=65)

open(unit=66, status = 'old',file = 'ephase.dat')
read (66,*)(ephase(k), k = 1,46)
write(6,*)' read ',k-1,' points from ephase'
close(unit=66)

WRITE(6,*)'ALL INPUT DATA READ'

> open(unit=67, status = 'unknown',file = 'Q_INV.RES')
> open(unit=70, status = 'unknown',file = 'qp.b7')
> open(unit=71, status = 'unknown',file = 'avg_ge.b7')
> open(unit=72, status = 'unknown',file = 'tp_qp.sim10')
> open(unit=73, status = 'unknown',file = 't_inv.sim10')
open(unit=74, status = 'unknown',file = 'avg_tamp.end')
open(unit=79, status = 'unknown',file = 'avg_tsamp.end')
> open(unit=80, status = 'unknown',file = 'tamp.sim10')
> open(unit=81, status = 'unknown',file = 't_diff.test1')

**KEYBOARD ENTERED VALUES**

WRITE(6,*)'ENTER VALUE FOR CLOUDINESS'
READ(5,*)CLOUD
WRITE(6,*)

WRITE(6,*)'ENTER NUMBER OF DEPTH NODES ,MUST BE EVEN'
READ(5,*)NODES
WRITE(6,*)

WRITE(6,*)'ENTER VALUE FOR THERMAL DIFFUSIVITY IN M2/SEC'
READ(5,*)DT
*CONVERT DT FROM M2/SEC TO CM2/MONTH
DT=DT*10000.*3600.*24.*365./12.
WRITE(6,*)

WRITE(6,*)'ENTER VALUE FOR NODE SPACING IN METERS'
READ(5,*)DELTAZ
*CONVERT FROM METERS TO CENTIMETERS
DELTAZ=DELTAZ*100.
WRITE(6,*)

WRITE(6,*)'ENTER VALUE FOR DELTA T IN YEARS'
READ(5,*)DELTAT
*CONVERT FROM YEARS TO MONTHS
DELTAT=12.*DELTAT
WRITE(6,*)

WRITE(6,*)'ENTER LATITUDE INCREMENTS IN DEGREES'
READ(5,*)PHIDINT
WRITE(6,*)

WRITE(6,*)'ENTER VALUE FOR HEAT CAPACITY OF EARTH IN CAL/CM3-C'

```

```

      READ(5,*)C
* CONVERT C FROM CAL/CM^3-C TO KCAL/CM^3-C
      C=.001*C
      WRITE(6,*)

      WRITE(6,*)'ENTER VALUE FOR GE TOLERANCE IN KCAL/CM^2 MONTH'
      READ(5,*)GETOL
      WRITE(6,*)

      WRITE(6,*)'ENTER VALUE FOR SEASONAL INCREMENT IN DEGREES'
      WRITE(6,*)'ONE MONTH EQUALS 30 DEGREES'
      READ(5,*)LAMDINT
      WRITE(6,*)

      WRITE(6,*)'ENTER VALUE FOR STARTING TIME'
      WRITE(6,*)'-10000. MEANS TEN THOUSAND YEARS AGO '
      WRITE(6,*)
      READ(5,*)START
      WRITE(6,*)

      WRITE(6,*)'ENTER VALUE FOR TEMPERATURE AMPLIFICATION'
      READ(5,*)AMP

      WRITE(6,*)'ENTER FINAL TIME'
      READ(5,*)TFINAL
**TIME LOOP**

      DO 40 TIME=START,TFINAL,DELTAT/12.

**CALCULATE AVERAGE YEARLY TEMPERATURE FOR THE EARTH**

      IF (TIME.EQ.START) THEN
        TP=6.718
        tpge=tp
      ELSE
        I=0
        SUMGE=0.
        SUMLATQ=0.
        SUMCOS=0.
        SUMTAMP=0.
        SUMTSAMP=0.
        DO 41 PHID=0.,90.,PHIDINT
          I=I+1
          phi = phiD*(pi/const2)
          SUMTAMP=SUMTAMP + TAMP(I)*COS(PHI)
          SUMTSAMP=SUMTSAMP + TSAMP(I)*COS(PHI)
          SUMGE=SUMGE + GE(I)*COS(PHI)
          SUMLATQ=SUMLATQ + Q(I)*COS(PHI)
          SUMCOS=SUMCOS + COS(PHI)
41      CONTINUE
        AVG_GE=SUMGE/SUMCOS
        QP=SUMLATQ/SUMCOS
        AVG_TAMP=SUMTAMP/SUMCOS
        AVG_TSAMP=SUMTSAMP/SUMCOS
c      IF (MOD(NINT(-TIME),100).EQ.0)WRITE(70,*)-time,QP
        TP=(QP*(1.-P_ALB)-A1+A2*CLOUD)/(BETA+B1-B2*CLOUD)
        TPGE=(QP*(1.-P_ALB)-AVG_GE-A1+A2*CLOUD)/(BETA+B1-B2*CLOUD)
      END IF

```

```

c      if(mod(-time,50.).eq.0.)WRITE(72,*)-TIME,tp,qp
      IF(MOD(-TIME,100.).EQ.0.)WRITE(6,*)-TIME,AVG_TAMP,AVG_TSAMP,
+      avg_ge
      if(mod(-time,100.).eq.0.)WRITE(74,*)-TIME,AVG_TAMP
      if(mod(-time,100.).eq.0.)WRITE(79,*)-TIME,AVG_TSAMP
      if(mod(-time,100.).eq.0.)WRITE(71,*)-TIME,AVG_GE

**LOOP FOR LATITUDES**
      I=0
      DO 50 PHID=0.,90.,PHIDINT
      I=I+1

**LOOP FOR SEASONAL VARIATIONS**

      COUNT=0.
      SUMQ=0.
      SUMWEIGHT=0.
      DO 60 lamD= 0.,(360.-LAMDINT),LAMDINT

**WEIGHTING WITH SEASON**

c      IF(LAMD.LT.90.)WEIGHT= .5 + LAMD/180.
c      IF(LAMD.GE.90.)WEIGHT= 1.5-LAMD/180.

      COUNT=COUNT+1.

**ALBEDO VALUES THAT DON'T VARY WITH TIME**

      IF(ABS(PHID).LE.(60.))THEN
        ALBEDO(I)=.32
      END IF
      IF((ABS(PHID).GT.(60.)).AND.(ABS(PHID).LE.(70.)))THEN
        ALBEDO(I)=.32+(phid-60.)*.018
      END IF
      IF(ABS(PHID).GT.(70.))THEN
        ALBEDO(I)=.50+(phid-70.)*.012
      END IF
      P_ALB=.334

** Convert to radians **

      phi = phid*(pi/const2)
      lam = lamD*(pi/const2)

C.....
      sumc1 = 0.
      sumc2 = 0.
      sumob = 0.

      do 200 j = 1,47

** Mid-month obliquity for a given LATITUDE AND TIME OF YEAR **

      +      sumob = sumob + (oamp(j)/const1)*(cos(((omnrt(j)*dble(time))
+      /const1 + ophase(j))*(pi/const2)))

** sumc1 & sumc2 are used to calculate eccentricity and longitude of the **
** perihelion based on a moving vernal equinox **

      sumc1 = sumc1 + eamp(j)*sin(((emnrt(j)*dble(time))/const1

```



```

+          + ephase(j))*(pi/const2))
+      sumc2 = sumc2 + eamp(j)*cos(((emnrt(j)*dble(time))/const1
+          + ephase(j))*(pi/const2))
?00      continue

      obliq = sumob + obconst
      long = (atan2(sumc1,sumc2))*(const2/pi)
      eccen = sumc1/sin(atan2(sumc1,sumc2))

** PARAMETERS USED TO CALCULATE MID-MONTH SOLAR INSOLATION **

      delta = asin((sin(lam))*sin(obliq*(pi/const2)))
      nu = lam-((long+const2)*(pi/const2))
      rho = (one-(eccen**sq))/
+          (one+(eccen*cos(nu)))

** MID-MONTH SOLAR INSOLATION IN CAL/CM^2-DAY **

      IF(ABS(PHI).GE.(PI/TWO-abs(DELTA)))THEN
+      if(((phi.gt.(0.)).and.(delta.gt.(0.))).or.((phi.lt.(0.)).
+          and.(delta.lt.(0.))))then
+      SOLRINS= ((insconst*solrconst)/(rho**sq))
+          *sin(phi)*sin(delta)
      end if
+      if(((phi.lt.(0.)).and.(delta.gt.(0.))).or.((phi.gt.(0.)).
+          and.(delta.lt.(0.))))then
+      SOLRINS = DBLE(0.)
      end if

      ELSE

      Ho = acos(-tan(phi)*tan(delta))

+      solrins = ((insconst*solrconst)/(pi*(rho)**sq))
+          *(Ho*sin(phi)*sin(delta)+cos(phi)*cos(delta)*sin(Ho))
      END IF

**CONVERT SOLAR INSOLATION VALUES FOR EACH LATITUDE**
**IN LATITUDE LOOP TO kcal/cm^2 month AND SUM VALUES OVER
**EACH LATITUDE

      SUMQ=SUMQ + SOLRINS/DBLE(32.88)
:      SUMWEIGHT=SUMWEIGHT+WEIGHT

60      CONTINUE

**CALCULATE AVERAGE YEARLY SOLAR INSOLATION VALUES FOR EACH
**LATITUDE
:      Q(I)=SUMQ/(COUNT)
:      Q(I)=SUMQ/(SUMWEIGHT)

```

```

> IF(MOD(NINT(-TIME),1000).EQ.0)WRITE(67,*)-TIME,Q(I)
**CALCULATE AVERAGE YEARLY TEMPERATURE FOR EACH LATITUDE INCREMENT
**USING AVERAGE YEARLY GE FROM PREVIOUS YEAR

+ TNOGE(I)=(Q(I)*(DBLE(1.)-ALBEDO(I))-A1+A2*CLOUD+BETA*TP)/
+ (BETA+B1-B2*CLOUD)

TAMP(I)=(TNOGE(I)-AVGT(I))*AMP + AVGT(I)

IF(ABS( GE(I)-LASTGE(I) ).LT.GETOL)THEN
TS(I)=(Q(I)*(DBLE(1.)-ALBEDO(I))-GE(I)-A1+
+ A2*CLOUD+BETA*TPGE)/(BETA+B1-B2*CLOUD)
ELSE
TS(I)=(Q(I)*(DBLE(1.)-ALBEDO(I))-A1+A2*CLOUD+BETA*TP)/
+ (BETA+B1-B2*CLOUD)
END IF

TSAMP(I)=(TS(I)-AVGT(I))*AMP + AVGT(I)

> WRITE(73,*)-TIME,TSAMP(I)
> WRITE(80,*)-TIME,TAMP(I)
> WRITE(81,*)-TIME,TAMP(I)-TSAMP(I)
> WRITE(72,*)-TIME,GE(I)

**INITIAL SUBSURFACE TEMPERATURES WITH INCLUSION OF GEOTHERMAL GRADIENT
**INITIALIZE PREVIOUS SUBSURFACE TEMPERATURES EQUAL TO PRESENT VALUES
**SET BOTTOM NODE TEMPERATURE TO REMAIN CONSTANT

IF(TIME.EQ.START)THEN
TEMP(I,1)=TSAMP(I)
TLAST(I,1)=TEMP(I,1)
DO 10 J=2,NODES
TEMP(I,J)=DBLE(J-1)*DELTAZ*DBLE(.0003) + TSAMP(I)
TLAST(I,J)=TEMP(I,J)
10 CONTINUE
END IF

**REASSIGN AVERAGE YEARLY SUBSURFACE TEMPERATURE TO EARLIER TIME**

DO 20 J=1,NODES
T2AGO(I,J)=TLAST(I,J)
TLAST(I,J)=TEMP(I,J)
20 CONTINUE

50 CONTINUE

**AFTER THE FIRST TIME ITERATION CAN CALCULATE CHANGE IN SUBSURFACE
**HEAT STORAGE , GE .

IF(TIME.GT.START)THEN
DO 17 I=1,7
LASTGE(I)=GE(I)
17 CONTINUE
CALL SUBSURF(NODES,C,GE,DELTAT,DELTAZ,T2AGO,TLAST,
+ TEMP,GRAD)
END IF

**CALCULATE SUBSURFACE TEMPERATURE DISTRIBUTIONS

```

CALL DEPTH(NODES-1,DELTAZ,deltat,TEMP,TLAST,DT,TSAMP)

```
40 CONTINUE
  3 write(72,*)
  3 write(73,*)
  write(74,*)
  write(71,*)
  3 write(70,*)
  3 write(67,*)
  write(79,*)
  3 write(80,*)
  3 write(81,*)
  3 CLOSE(UNIT=67)
  3 CLOSE(UNIT=70)
  3 CLOSE(UNIT=71)
  3 CLOSE(UNIT=72)
  3 CLOSE(UNIT=73)
  CLOSE(UNIT=74)
  CLOSE(UNIT=79)
  3 CLOSE(UNIT=80)
  3 CLOSE(UNIT=81)
  STOP
  END
```

```
**DEPTH*****
* SUBROUTINE USING *
* IMPLICIT 1-D TRANSIENT FINITE DIFFERENCE CODE TO FIND TEMPERATURE *
* AT NODES BELOW SURFACE OF EARTH . *
* SOLVES PARABOLIC PARTIAL DIFFERENTIAL EQUATION USING CRANK- *
* NICOLSON ALGORITHM. *
* *
* NODES : INTEGER, THE NUMBER OF DEPTH NODES . [-] *
* TEMP(NODES): TEMPERTURE AT NODE . [C] *
* DELTAZ: DISTANCE BETWEEN NODES . [CM] *
* DELTAT: TIME INCREMENT . [MONTH] *
* DT: THERMAL DIFFUSIVITY . [CM^2/MONTH] *
* TLAST(NODES): TEMPERATURE AT NODES THE PREVIOUS TIME INCREMENT.[C] *
* TSURF: TEMPERATURE AT SURFACE NODE FOR THIS TIME INCREMENT.[C] *
* TBOT : CONSTANT TEMPERATURE AT BOTTOM NODE . [C] *
*****
```

SUBROUTINE DEPTH(M,H,deltat,TEMP,TLAST,DT,TSURF)

```
DOUBLE PRECISION TEMP(1:7,0:M),H,deltat,DT,
+ TSURF(7),TLAST(1:7,0:M),LAMDA,A(1000,1000),
+ XO(7,1000),B(1000),X(1000)
```

```
TOTALD=H*REAL(M)
LAMDA=DT*deltat/(H*H)
TOL=.000001
OMEGA=DBLE(1.24)
```

```
DO 3 ILAT=1,7
  DO 4 I=0,M
    XO(ILAT,I)=TLAST(ILAT,I)
    X(I)=0.
```

```
4 CONTINUE
3 CONTINUE
```

```
DO 60 ILAT=1,7
```

```

+      B(1)=(1.-LAMDA)*TLAST(ILAT,1)+(LAMDA/2.)*(TLAST(ILAT,2)
+          +TLAST(ILAT,0)+TSURF(ILAT))
      B(M)=.0003*H
      A(1,1)=1.+LAMDA
      A(1,2)=-LAMDA/2.
      A(M,M-1)=-1.
      A(M,M)=1.

      DO 5 I=2,M-1
+          B(I)=(1.-LAMDA)*TLAST(ILAT,I)+(LAMDA/2.)*
+              (TLAST(ILAT,I-1)+TLAST(ILAT,I+1))
+          A(I,I-1)=-LAMDA/2.
+          A(I,I)=1.+LAMDA
+          A(I,I+1)=-LAMDA/2.
5      CONTINUE

      DO 7 L=1,1000
          BIGDIF=0.

          DO 6 I=1,M
+              X(I)=(1.-OMEGA)*XO(ILAT,I)+OMEGA*( B(I)-A(I,I-1)*X(I-1)
+                  -A(I,I+1)*XO(ILAT,I+1) )/A(I,I)

+              IF(ABS(X(I)-XO(ILAT,I)).GT.BIGDIF)THEN
+                  BIGDIF=ABS(X(I)-XO(ILAT,I))
+              END IF
6          CONTINUE
          IF(BIGDIF.LT.TOL)GO TO 50

          DO 8 I=1,M
+              XO(ILAT,I)=X(I)
8          CONTINUE

7          CONTINUE
          WRITE(6,*)'EXCEEDED MAXIMUM ITERATIONS'

50         CONTINUE

          DO 500 I=1,M
+              TEMP(ILAT,I)=X(I)
500        CONTINUE

60        CONTINUE

      END

```

```

*SUBSURF*****
* SUBROUTINE TO DETERMINE CHANGE IN HEAT STORAGE OF EARTH OVER ONE *
* TIME INTERVAL . DERIVATIVES ARE FOUND USING THREE TEMPERTURE *
* POINTS AND LAGRANGE COEFFICIENT POLYNOMIALS . THE INTEGRATION IS *
* PREFORMED USING A SIMPSON COMPOSITE NUMERICAL INTEGRATION SCHEME . *
* *
* C :HEAT CAPACITY OF EARTH . [KCAL/CM3-K] *
* TOTALD : TOTAL DEPTH FROM SURFACE TO LOWER CONSTANT TEMPERATURE *
* BOUNDARY . [CM] *
* GE :CHANGE IN HEAT STORAGE OF EARTH . [KCAL/CM2-MONTH] *
* GRAD : TEMPERATURE GRADIENT WRT TIME FOR EACH NODE . [C/MONTH] *
*****

```

```

SUBROUTINE SUBSURF(NODES,C,GE,DELTAT,DELTAZ,T2AGO,TLAST,

```

```

+           TEMP, GRAD)

DOUBLE PRECISION C, GE(7), DELTAT, DELTAZ, T2AGO(7, NODES),
+         TLAST(7, NODES), TEMP(7, NODES), GRAD(7, NODES), TOTALD,
+         ZI0, ZI1, ZI2, ENDA, ENDB, H

INTEGER NODES

TOTALD=DBLE(NODES-1)*DELTAZ

****FIND TEMPERATURE GRADIENT WRT TIME*****
*FIRST NODE (SURFACE TEMPERATURE) GRADIENTS ARE NOT TAKEN INTO ACCOUNT
*****

      DO 11 I=1,7
      DO 10 J=2, NODES
          GRAD(I, J)=(-DBLE(1.)/DELTAT)*(-DBLE(1.5)*TEMP(I, J)+
+          DBLE(2.) *TLAST(I, J)-DBLE(.5)*T2AGO(I, J))
10      CONTINUE
11      CONTINUE

****INTEGRATE TEMPERATURE GRADIENT WRT DEPTH*****

      ENDA=DELTAZ
      ENDB=TOTALD
      M=(NODES-2)/2
      H=(ENDB-ENDA)/DBLE(2*M)
      DO 21 I=1,7
          ZI0=GRAD(I, 2)+GRAD(I, NODES)
          ZI1=DBLE(0.)
          ZI2=DBLE(0.)
          DO 20 J=2, NODES-2
              IF(MOD(J, 2).EQ.0) ZI2=ZI2+GRAD(I, J+1)
              IF(MOD(J, 2).NE.0) ZI1=ZI1+GRAD(I, J+1)
20          CONTINUE
          GE(I)=C*H*(ZI0 +DBLE(2.)*ZI2 + DBLE(4.)*ZI1)/DBLE(3.)
21          CONTINUE
      END

```

APPENDIX III

```

PROGRAM PLATE.F
      CREATE A DATA FILE TO USE FOR A SPECTRAL CONTOUR PLOT SIMILAR
      TO THE ONE SEEN IN PALEOCEANOGRAPHY,AUGUST 1990,PG.519.
.....
DIMENSION T(10000),X(10000),TB(1000),XB(1000)

CHARACTER*20INFILE,OUTFILE1,OUTFILE2

WRITE(6,*)'ENTER NAME OF INPUT FILE'
READ(5,'(A)')INFILE
WRITE(6,*)'ENTER NAME OF OUTPUT FILE 1 , SPECTUM DATA'
WRITE(6,*)
READ(5,'(A)')OUTFILE1
WRITE(6,*)'ENTER NAME OF OUTPUT FILE 2 , MEAN AND VARIANCE'
WRITE(6,*)
READ(5,'(A)')OUTFILE2

OPEN(UNIT=21,FILE=INFILE,STATUS='OLD')
OPEN(UNIT=22,FILE=OUTFILE1,STATUS='UNKNOWN')
OPEN(UNIT=23,FILE=OUTFILE2,STATUS='UNKNOWN')
.....
      READ IN DATA
.....
      WRITE(6,*)'ENTER TOTAL # OF SPECTRUMS DESIRED <27>'
      READ(5,*)NSPECT
      WRITE(6,*)'ENTER INTEGER # FOR BAND WIDTH <200>'
      READ(5,*)NBAND
      WRITE(6,*)'ENTER LENGTH OF SHIFT INTERVAL IN KA'
      READ(5,*)SHIFT
      WRITE(6,*)'ENTER STARTING POINT'
      READ(5,*)START

DO 50 I=1,10000
      READ(21,*,END=100)T(I),X(I)
      NCOUNT=NCOUNT + 1
50  CONTINUE
100 CONTINUE
      BAND=NBAND

      DO 20 J=1,NSPECT
.....
      FIND MEAN FOR INTERVAL
.....
      TMAX=START-(J-1)*SHIFT
      TMIN=TMAX-BAND
      WRITE(6,*)'TMAX=',TMAX,' TMIN=',TMIN
      IA=0
      DO 105 I=1,NCOUNT
      IF((T(I).LE.TMAX).AND.(T(I).GE.TMIN))THEN
      IA=IA+1
      TB(IA)=T(I)
      XB(IA)=X(I)
      ELSE
      END IF
105  CONTINUE
      SUM=0.
      DO 106 I=1,IA
      SUM=SUM+XB(I)
106  CONTINUE
      XIA=IA
      XMEAN=SUM/XIA
      WRITE(6,*)'MEAN=',XMEAN
      -----

```


DIMENSION X(NCOUNT), T(NCOUNT)

REALPI = 3.1414926536

TOP1 = 0.

BOT1 = 0.

TOP2 = 0.

BOT2 = 0.

DO 100 I=1, NCOUNT

TERM = 2.* REALPI * FREQ * (T(I) - TAU)

TOP1 = TOP1 + (X(I) - XMEAN) * COS(TERM)

BOT1 = BOT1 + COS(TERM) * COS(TERM)

TOP2 = TOP2 + (X(I) - XMEAN) * SIN(TERM)

BOT2 = BOT2 + SIN(TERM) * SIN(TERM)

100 CONTINUE

QUANT = TOP1 * TOP1 / BOT1 + TOP2 * TOP2 / BOT2

PER = QUANT / (2. * VARI)

END